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Geometric inequalities for spacelike hypersurfaces in the Minkowski spacetime $\stackrel{\text{the}}{\Rightarrow}$

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Abstract

We derive a linear isoperimetric inequality and some geometric inequalities for properly located compact achronal spacelike hypersurfaces via a Minkowski-type integral formula in the Minkowski spacetime. © 2001 Elsevier Science B.V. All rights reserved.

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1. Introduction

Let *M* be a compact achronal spacelike hypersurface in the Minkowski (n + 1)-spacetime \mathbb{L}^{n+1} contained in the chronological future $I^+(O)$ of the origin *O* in \mathbb{L}^{n+1} . Let \mathcal{K}_M denote the *cone* of *M* (with respect to *O*) given by

$$\mathcal{K}_M = \{ \lambda p \in \mathbb{L}^{n+1} : p \in M, 0 \le \lambda \le 1 \},\$$

and $\Omega = \Omega(M)$ the normalized hypersurface of M given by

$$\Omega = \Omega(M) = \left\{ \frac{x}{\|x\|} \in \mathbb{L}^{n+1} : x \in M \right\},\$$

where ||x|| > 0 is the norm of x. Note that $\Omega(M)$ is also a compact achronal spacelike hypersurface contained in $I^+(O)$. By V(B), we denote the (n + 1)-dimensional volume of $B \subset \mathbb{L}^{n+1}$ and by A(B) the *n*-dimensional (Lorentzian) volume of $B \subset \mathbb{L}^{n+1}$. Recently,

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Bahn and Ehrlich [3] obtained the following isoperimetric inequality for such a spacelike hypersurface in the Minkowski spacetime \mathbb{L}^{n+1} as a consequence of a Brunn–Minkowski type inequality in the Minkowski spacetime \mathbb{L}^{n+1} .

Theorem 1. Let *M* be a compact achronal spacelike hypersurface in \mathbb{L}^{n+1} contained in $I^+(O)$ with the (Lorentzian) distance $d(O, M) = t^*$. Then

$$A^{n+1}(M) \le (n+1)^{n+1} V(\mathcal{K}_{\Omega}) V^n(\mathcal{K}_M)$$

where equality holds only when M is contained in the (upper) hyperbolic space $\mathbb{H}^{n}(t^{*})$ of radius t^{*} .

This isoperimetric inequality is a Lorentzian version of the isoperimetric inequality for convex cones of Lions and Pacella [7].

In this paper, we consider geometric inequalities for a compact spacelike hypersurface M in the Minkowski (n+1)-spacetime \mathbb{L}^{n+1} involving the mean curvature H of M. Notice that the quantities $V(\mathcal{K}_{\Omega})$ and $V(\mathcal{K}_M)$ in Theorem 1 depend on the location of M in \mathbb{L}^{n+1} . So, for our purpose, we restrict our attention to a compact spacelike hypersurface $M \subset \mathbb{L}^{n+1}$ such that M is contained in the chronological future $I^+(O)$ of the origin O in \mathbb{L}^{n+1} and

$$\int_{\partial M} \langle P, \nu \rangle \, \mathrm{d}S \ge 0,\tag{1}$$

where ν is the outward unit conormal vector field along ∂M and P the position vector field in \mathbb{L}^{n+1} . Such a compact spacelike hypersurface M will be said to be *properly located*. Note that " $\langle P, \nu \rangle \ge 0$ " represents that the hyperbolic angle from ν to P is non-negative (cf. [2,4]). For some properly located compact spacelike hypersurfaces $M \subset \mathbb{L}^{n+1}$, we derive a linear isoperimetric inequality via a Minkowski-type integral formula (see Lemma 3). Finally, we prove the following geometric inequalities in Theorems 6 and 7.

Theorem 2. Let *M* be a properly located compact achronal spacelike hypersurface in \mathbb{L}^{n+1} with mean curvature $H \ge c > 0$. Then

$$V(\mathcal{K}_M) \le \frac{1}{c^{n+1}} V(\mathcal{K}_{\Omega}), \qquad A(M) \le \frac{1}{c^n} A(\Omega)$$

with equality only when M is a subset of $\mathbb{H}^n(1/c)$.

Note that, for our convention, the (upper) hyperbolic space $\mathbb{H}^{n+1}(r)$ of radius r > 0 has positive constant mean curvature 1/r.

For detailed account of the terminology relating to indefinite metric and causality, the reader may consult [5] or [8].

2. Minkowski-type integral formula

The Minkowski (n + 1)-spacetime \mathbb{L}^{n+1} $(n \ge 2)$ is just the real vector space \mathbb{R}^{n+1} with the scalar product *g* of index 1,

$$g(x, y) = \langle x, y \rangle = -x_0 y_0 + \sum_{i=1}^n x_i y_i$$

for $x = (x_0, x_1, ..., x_n)$, $y = (y_0, y_1, ..., y_n) \in \mathbb{R}^{n+1}$. We will assume, as usual, \mathbb{L}^{n+1} is time-oriented by the unit timelike vector $e_0 = (1, 0, ..., 0)$ so that a non-spacelike tangent vector $v \in \mathbb{L}^{n+1}$ is said to be future-directed if $\langle e_0, v \rangle < 0$. We denote by $\mathbb{H}^n(r)$ the (*upper*) *hyperbolic n-space* of radius r > 0 in \mathbb{L}^{n+1} given by

$$\mathbb{H}^{n}(r) = \{ x = (x_0, x_1, \dots, x_n) \in \mathbb{L}^{n+1} : \langle x, x \rangle = -r^2, x_0 > 0 \}$$

which is a complete spacelike hypersurface of in the Minkowski spacetime \mathbb{L}^{n+1} with constant curvature $-1/r^2$.

Let *M* be a compact spacelike hypersurface in \mathbb{L}^{n+1} oriented by the future-directed unit timelike normal vector field *N*. Note that every compact spacelike hypersurface *M* in \mathbb{L}^{n+1} necessarily has a non-empty boundary ∂M (cf. [1]). So, for simplicity, we will always assume that *M* has smooth boundary ∂M . We will decompose the position vector field *P*, restricted on *M*, as

$$P = P^{\mathrm{T}} - \langle P, N \rangle N, \tag{2}$$

where P^{T} denotes the tangential part of P to M. Let $\overline{\nabla}$ and ∇ denote the Levi–Civita connections of \mathbb{L}^{n+1} and M, respectively. Then, the Gauss and Weingarten formulas for M in \mathbb{L}^{n+1} are written, respectively, as

$$\nabla_X Y = \nabla_X Y - \langle S(X), Y \rangle N, \tag{3}$$

$$S(X) = -\nabla_X N,\tag{4}$$

for all smooth tangent vector fields $X, Y \in \mathfrak{X}(M)$ on M, where S is the shape operator of M in \mathbb{L}^{n+1} associated with N. The mean curvature H on M is given by

$$H = -\frac{1}{n}\operatorname{tr}(S). \tag{5}$$

Now, we can formulate a Lorentzian version of a Minkowski formula (cf. [6]) in the Minkowski spacetime as given in the following lemma.

Lemma 3. Let M be a compact spacelike hypersurface with smooth boundary ∂M in the Minkowski spacetime \mathbb{L}^{n+1} . Then

$$\frac{1}{n} \oint_{\partial M} \langle P, \nu \rangle \, \mathrm{d}S = A(M) + \int_{M} \langle P, N \rangle H \, \mathrm{d}A, \tag{6}$$

where dA is the n-dimensional volume element on M, dS the induced (n - 1)-dimensional volume element on ∂M .

Proof. Let $\phi = \langle P, P \rangle$, where *P* is the position vector field in \mathbb{L}^{n+1} . Then, the gradient of ϕ on *M* is given by

$$\nabla \phi = 2P^{\mathrm{T}} = 2(P + \langle P, N \rangle N),$$

so the Hessian of ϕ on M is given as

$$\nabla^2 \phi(X, Y) = \langle \nabla_X (\nabla \phi), Y \rangle = 2 \langle X, Y \rangle - 2 \langle P, N \rangle \langle S(X), Y \rangle,$$

where $X, Y \in \mathfrak{X}(M)$ by (3) and (4). Therefore, by (5), the Laplacian of ϕ on M is given by

$$\Delta \phi = 2n + 2n \langle P, N \rangle H.$$

Integrating this over M, by the divergence theorem, we have

$$\oint_{\partial M} \langle P^{\mathrm{T}}, \nu \rangle \,\mathrm{d}S = nA(M) + n \int_{M} \langle P, N \rangle H \,\mathrm{d}A$$

which implies (6).

Remark 4. Let *M* be a compact spacelike hypersurface in \mathbb{L}^{n+1} with smooth boundary ∂M . Note that the condition (1) is equivalent to the weak super-harmonicity of $\phi = \langle P, P \rangle$, i.e.,

$$\int_M \Delta \phi \, \mathrm{d}A \ge 0.$$

Let $\Pi_p(P, \nu)$ be the subspace of tangent space $T_p(\mathbb{L}^{n+1})$ of \mathbb{L}^{n+1} at $p \in \partial M$ spanned by P_p and ν_p . Let $\{u_0, u_1\}$ be an orthonormal basis of $\Pi_p(P, \nu)$, where u_0 is a future-directed unit timelike vector and $\langle u_1, \nu_p \rangle > 0$. Then, the scalar product of P_p and ν_p is given by

$$\langle P_p, v_p \rangle = ||P_p|| \sinh(\varphi(P_p) - \varphi(v_p)),$$

where $\varphi(v)$ is the *hyperbolic angle* of v with respect to $\{u_0, u_1\}$. Notice that the hyperbolic angle from v_p to P_p , " $\varphi(P_p) - \varphi(v_p)$ ", is independent on the choice of the orthonormal basis $\{u_0, u_1\}$ in $\Pi_p(P, v)$ (cf. [2,4]). In particular, for a compact spacelike hypersurface in \mathbb{L}^{n+1} with smooth boundary $\partial M \subset \mathbb{H}^n(r)$, the condition (1) means that the hyperbolic angle $\varphi(P)$ is greater than or equal to the hyperbolic angle $\varphi(v)$ in $\Pi(P, v)$ on average. For instance, consider the family of hyperbolic caps

$$M_c = \{ x \in \mathbb{L}^{n+1} : \langle x, x \rangle = -c^2, 0 < x_0 \le \sqrt{1+c^2} \} \quad (0 < c < \infty)$$

Then all M_c are properly located, since $M_c \subset I^+(O)$ and $\langle P, \nu \rangle = 0$ on the boundary ∂M_c . Furthermore,

$$M_c - \vec{\varepsilon} = \{p - \vec{\varepsilon} \in \mathbb{L}^{n+1} : p \in M_c, \quad \vec{\varepsilon} = (\varepsilon, 0, \dots, 0), \varepsilon \ge 0\}$$

is also properly located if $M_c - \vec{\varepsilon}$ is still contained in $I^+(O)$.

3. Geometric inequalities

We first derive a linear isoperimetric inequality for a properly located compact achronal spacelike hypersurface, which is given in the following lemma.

Lemma 5. Let *M* be a properly located compact achronal spacelike hypersurface in \mathbb{L}^{n+1} with mean curvature $H \ge c > 0$. Then

$$V(\mathcal{K}_M) \le \frac{1}{c(n+1)} A(M).$$
⁽⁷⁾

Proof. Since *M* is properly located, from the Minkowski formula (Lemma 3),

$$-\int_M \langle P, N \rangle H \, \mathrm{d}A \le A(M)$$

Note that *P* and *N* are future-directed, so $\langle P, N \rangle < 0$ on *M*. From the condition that $H \ge c > 0$, we have

$$-c \int_M \langle P, N \rangle \, \mathrm{d}A \le A(M).$$

Note that the (n + 1)-dimensional volume of the cone \mathcal{K}_M of M is

$$V(\mathcal{K}_M) = -\frac{1}{n+1} \int_M \langle P, N \rangle \, \mathrm{d}A.$$

Thus, we have

$$c(n+1)V(\mathcal{K}_M) \le A(M),$$

which is equivalent to (7).

Theorem 6. Let *M* be a properly located compact achronal spacelike hypersurface in \mathbb{L}^{n+1} with mean curvature $H \ge c > 0$. Then

$$V(\mathcal{K}_M) \le \frac{1}{c^{n+1}} V(\mathcal{K}_{\Omega}) \tag{8}$$

with equality only when M is a subset of $\mathbb{H}^n(1/c)$.

Proof. From Theorem 1 and Lemma 5, we have

$$(n+1)cV(\mathcal{K}_M) \le (n+1)V^{1/(n+1)}(\mathcal{K}_{\Omega})V^{n/(n+1)}(\mathcal{K}_M),$$

which is equivalent to (8). Suppose that $M \subset \mathbb{H}^n(1/c)$. Then

$$\mathcal{K}_M = \frac{1}{c} \mathcal{K}_{\Omega} = \left\{ \frac{1}{c} q \in \mathbb{L}^{n+1} : q \in \mathcal{K}_{\Omega} \right\}.$$

Thus, the equality in (8) clearly holds. Conversely, suppose that

$$V(\mathcal{K}_M) = \frac{1}{c^{n+1}} V(\mathcal{K}_{\Omega}).$$
(9)

Then, by Lemma 5, we have

 $A^{n+1}(M) \le (n+1)^{n+1} V(\mathcal{K}_{\Omega}) V^n(\mathcal{K}_M).$

By Theorem 1, we have $M \subset \mathbb{H}^n(t^*)$, where $t^* = d(O, M)$, the (Lorentzian) distance from the origin O to M (cf. [2,4]). By (9), we see $t^* = 1/c$ and so $M \subset \mathbb{H}^n(1/c)$.

Theorem 7. Let *M* be a properly located compact achronal spacelike hypersurface in \mathbb{L}^{n+1} with mean curvature $H \ge c > 0$. Then

$$A(M) \le \frac{1}{c^n} A(\Omega) \tag{10}$$

with equality only when M is a subset of $\mathbb{H}^n(1/c)$.

Proof. From Theorems 1 and 6, we have

$$A(M) \le \frac{1}{c^n}(n+1)V(\mathcal{K}_{\Omega})$$

Since $\Omega \subset \mathbb{H}^n(1)$,

$$A(\Omega) = (n+1)V(\mathcal{K}_{\Omega}).$$

Thus, we get (10). Suppose that $M \subset \mathbb{H}^n(1/c)$. Then, the equality in (10) clearly holds. Conversely, suppose that

$$A(M) = \frac{1}{c^n} A(\Omega).$$
⁽¹¹⁾

Then, by Theorem 1, we have

$$\left(\frac{1}{c^{n+1}}\right)^n V^n(\mathcal{K}_{\Omega}) \le V^n(\mathcal{K}_M).$$

By Theorem 6, we have $M \subset \mathbb{H}^n(1/c)$.

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